

Overcoming surface energy to control Cu₃N epitaxial growth

Zainab Fatima¹, Isao Ohkubo^{1,3,*}, Satoshi Ishii^{1,3}, Takahiro Nagata², Takashi Aizawa¹, Takao Mori^{1,3}

¹Research Center for Materials Nanoarchitectonics (MANA), National Institute for Materials Science (NIMS), 1-1 Namiki, Tsukuba, Ibaraki 305-0044, Japan

²Research Center for Electronic and Optical Materials, National Institute for Materials Science (NIMS), 1-1 Namiki, Tsukuba, Ibaraki 305-0044, Japan

³Department of Materials Science, Institute of Pure and Applied Sciences, University of Tsukuba, 1-1-1 Tennodai, Tsukuba, Ibaraki 305-8573, Japan

* Corresponding author: OHKUBO.Isao@nims.go.jp

ABSTRACT

Crystal orientation control during copper nitride (Cu₃N) epitaxial growth was achieved using reactive DC magnetron sputtering. Both the (100)-orientation and (111)-orientation were observed readily from X-ray diffraction measurements for Cu₃N thin films grown on single-crystal MgO(100), MgO(111), SrTiO₃(100), and sapphire(0001) substrates. The Cu₃N(111) surface energy is greater than that of Cu₃N(100), suggesting that the Cu₃N(111) orientation has a lower formation probability than the Cu₃N(100) orientation. To control the influence of surface energy, thin film growth parameters related to the thermodynamics and kinetics of epitaxial thin film growth were tuned. Growth of single (111)-oriented Cu₃N epitaxial thin films, which has a higher surface energy orientation, was achieved on MgO(111) substrates. Optical band gaps of the single (111)-oriented Cu₃N epitaxial thin film were 1.80 eV for direct transition and 0.82 eV for indirect transition, indicating formation of a reasonable electronic structure in single (111)-oriented Cu₃N epitaxial thin films with higher surface energy.

1. INTRODUCTION

Thin film growth is a crystal growth process by which a vapor-phase species adsorbs onto a single-crystal substrate, leading to formation of a crystal lattice aligned under the influence of the substrate surface lattice. This process is driven both thermodynamically and kinetically because the system tends toward minimizing its energy while enabling atomic rearrangement through surface mobility.¹ Production of crystal-oriented thin films, i.e., epitaxial thin films, requires that a fine balance be achieved among chemical reactivity, thermal activation, and crystallographic lattice matching at the film-substrate interface.¹ For many research studies today, phase-pure and orientation-pure epitaxial thin films are necessary. During the initial growth stage of thin film formation, the surface energies and the nucleation process on a substrate can be expected to influence the crystal orientation of a thin film.²⁻⁴ Moreover, crystal orientation stability includes surface energy effects. Therefore, to achieve phase-pure and orientation-pure epitaxial thin film growth, the crystal orientation stability should be controlled by tuning the thin film epitaxial growth parameters associated with thermodynamics and kinetics.³⁻⁶

In this regard, copper nitride (Cu_3N), which is a particularly intriguing material with a cubic anti- ReO_3 structure, has attracted attention for various applications.⁷⁻⁹ The thin film growth of Cu_3N has been investigated using various growth techniques.¹⁰⁻²⁷ Actually, Cu_3N is prone to decomposition into Cu and N_2 , necessitating careful control of process parameters such as the growth temperature and nitrogen partial pressure, to stabilize the desired phase.¹² Among various thin film growth techniques, sputtering is more controlled in terms of kinetics than either molecular beam epitaxy (MBE) or chemical vapor deposition (CVD) because of the high kinetic energies of supply species from the sputtering target.²⁸ Therefore, process parameters such as the sputtering power, gas pressure, and growth rate influence the energy and supply species, leading to thin film formation. For example, high growth rates might favor amorphous or polycrystalline films because of limited surface diffusion and growth temperature, whereas optimized growth rates and substrate heating can promote crystallinity and preferred orientations.

Surface energy is another parameter that must be considered when elucidating the growth dynamics.²⁻⁴ Preferential growth of certain crystallographic orientations (100) or (111) textured domains is influenced by minimization of the surface and interfacial energies. For instance, low-energy surfaces predominate under normal conditions, but kinetic limitations (e.g., limited surface species mobility) can allow them to generate higher-energy surfaces. Additionally, differences in surface energy between the film and the substrate can lead to island formation, coalescence, and overall film roughness.²⁹

For this study, the growth of Cu₃N thin films by reactive DC magnetron sputtering is assessed on various single-crystal substrates, with specific examination of how thermodynamics, kinetics, and surface energy determine the Cu₃N thin films' crystallographic orientation. By varying growth parameters such as the growth temperature, growth rate, nitrogen gas pressure, and sputtering power, we can elucidate their roles in stabilizing the Cu₃N phase when controlling crystallographic orientation. These insights not only advance the fundamental understanding of nitride thin film synthesis: they are also necessary for achieving epitaxial growth and for tailoring film properties.

2. EXPERIMENTAL

The Cu₃N thin films used for this study were grown using reactive DC magnetron sputtering equipment (Kenix Co. Ltd.) with a 99.99% pure Cu target. Single-crystal sapphire(0001) (Cryscore Optoelectronic Ltd.), SrTiO₃(100) (Shinkosha Co. Ltd.), MgO(111) and MgO(100) substrates produced by CrysTec GmbH were used to assess thin film growth. All the substrates were preheated in a vacuum for 10–30 min using a lamp heater at 800 °C. The typical base pressure of the sputtering chamber was maintained on the order of 10⁻⁷ Pa. The DC power was 25–50 W. The sputtering gas was 99.999% pure nitrogen in the optimum range of 0.3–2.3 Pa. Its flow rate was 3–10 sccm. The separation distance between the target and substrate was 60–70 mm. Cu₃N thin films were grown at temperatures of 110–200 °C with growth rates of 0.15–1.0 Å/s because Cu₃N decomposes at 260 °C.^{10,17} Structural evaluations of Cu₃N thin films were performed based on X-ray diffraction

(XRD) measurements using a four-axis diffractometer (D8; Bruker Analytik GmbH) with Cu $K_{\alpha 1}$ radiation and a high-resolution two-dimensional (2D) detector. The 2D detector system can detect parts of the Debye–Scherrer ring rapidly and two-dimensionally with a psi (ψ) angle range of about 30° . The conventional 2θ – ω chart is obtained by integrating the 2D scan image in the ψ direction. The film thicknesses were measured using a surface profilometer (Dektak XT; Bruker Analytik GmbH). Chemical depth analysis of Cu_3N thin films was performed using ex-situ Auger electron spectroscopy (AES). The surface characterization apparatus, which had base pressure in the 10^{-8} Pa range, was equipped with a cylindrical mirror analyzer for AES and an ion gun for sputter etching.^{30,31} For Auger electron excitation, the 15 keV electron beam was focused on the normal surface of Cu_3N thin films at an angle of incidence of 70° . Sputter etching was performed with 3 keV Kr^+ ions at about 5 μA and with the angle of incidence of 45° . Surface morphologies of Cu_3N thin films were characterized using atomic force microscopy (AFM, AFM5300E; Hitachi High-Technologies Corp.). Optical band gaps of epitaxial Cu_3N thin films were characterized based on their absorptivity, obtained through spectroscopic ellipsometry measurements (SE 850; Sentech Instruments GmbH) at room temperature. The Tauc–Lorentz model³² was used to fit the ellipsometry spectra. Electronic structure calculations were performed using density-functional theory (DFT) through a full-potential linearized augmented plane-wave approach implemented in the WIEN2k code.³³ The modified Becke–Johnson (mBJ) potential^{34, 35} was employed. Energy band gaps (E_g) calculated using the mBJ potential are much closer to the experimentally obtained values than those from the generalized gradient approximation (GGA) functional potential and local density approximation (LDA).^{34–36} Parameter $R_{\text{mt}}K_{\text{max}}$, where R_{mt} represents the smallest muffin-tin radius, and where K_{max} denotes the plane-wave cutoff, was used to control the basis set size: this parameter was set to the high value of 7.0. For the atomic spheres, the R_{mt} values were 1.97 for Cu and 1.61 for N. In addition, to calculate the electronic structure, 100 k-points in the Brillouin zone were used. Convergence was reached with k-point sampling of 100. For the DFT calculations, we used

structural coordinates and lattice parameters⁹ of Cu₃N obtained from X-ray diffraction measurements.

3. RESULTS and DISCUSSION

Figure 1 shows typical XRD 2D images and 2θ - ω patterns of Cu₃N thin films grown on single-crystal sapphire(0001), SrTiO₃(100), MgO(111), and MgO(100) substrates at growth temperatures of 170–180 °C. On all substrates, multiple XRD diffraction peaks were observed with various crystallographic orientations originating from the Cu₃N phase, indicating that the thin films are Cu₃N polycrystalline. Whereas (100)-oriented growth of Cu₃N is typically favored on cubic substrates such as MgO(100) and SrTiO₃(100), strong (111) diffraction peaks were observed in both cases. By contrast, on sapphire(0001) substrates, where (111) orientation of Cu₃N was anticipated because of rhombohedral symmetry, dominant (100) diffraction peaks were observed.¹¹ For the MgO(111) substrates, (111) orientation peaks of Cu₃N were dominant, but were accompanied by other orientation peaks. These behaviors are more or less consistent with most previously reported Cu₃N thin film growth on glass and various single-crystal substrates, where peaks of (111)-oriented Cu₃N tend to be observed in addition to a peak of (100)-oriented Cu₃N, irrespective of the thin film growth method which was used.^{13,15,19,23–26} However, Cu₃N thin films produced using pulsed laser deposition show the (111) preferential growth of Cu₃N on Si(111) substrates.¹⁴ This growth suggests that the preferential growth orientation of Cu₃N depends on the growth methods and conditions. The lattice constant values of Cu₃N thin films obtained from XRD patterns presented in Fig. 1 were 3.828 Å for the (100) orientation and 3.826 Å for the (111) orientation on MgO(100) substrate, 3.820 Å for the (111) orientation on SrTiO₃(100) substrate, 3.827 Å for both (100) and (111) orientations on sapphire(0001) substrate, 3.822 Å for (100) orientation and 3.827 Å for (111) orientation on MgO(111) substrate. Comparison of these values to the bulk lattice constant of 3.819 Å⁹ shows that the lattice constant differences are 0.03–0.24%, suggesting that the Cu and N ratios might be approximately equal to the stoichiometric composition. Auger electron spectra of Cu₃N thin films

in the depth direction are portrayed in Fig. S1. In addition to Cu peaks, very weak carbon and oxygen peaks were observed. As portrayed in an AFM image of the Cu₃N thin film surface presented in Fig. S2, the origins of these carbon and oxygen peaks are regarded as residual carbon and oxygen in the depressions of Cu₃N thin film surfaces. Details of AES spectra are presented in the supplementary material. The influences of oxygen and carbon impurities are sufficiently small in this study.

For thin film growth, surface energy should play a central role in dictating the morphology, crystalline orientation, and overall structural quality of thin films.²⁻⁴ In fact, surface energy is associated with atoms near the surface because of their under-coordinated bonding environment compared to the bulk. For this study, surface energy ($Y_{(hkl)}$) was calculated using the equation derived by Holzer, Tielemann, and Müller *et al.*^{37,38} as shown below.

$$Y_{(hkl)} = \frac{1}{N_a} \Sigma \frac{S_{(hkl)} D^\circ}{A_{(hkl)}} \quad (1)$$

Therein, N_a is Avogadro's constant, $S_{(hkl)}$ represents the number of broken metal cation–anion bonds in $A_{(hkl)}$, where $A_{(hkl)}$ stands for the (hkl) surface area of the surface unit cell. In addition, D° denotes the bond energy. Actually, $A_{(hkl)}$ is calculated from the crystal structure data^{9,39-41} presented in Table 1. For this study, the D° values were calculated from the formation enthalpies (ΔH_f°) of crystal and gas phases for metal cation (M), anion (X), and a solid compound (MX) using the thermochemical equation^{42,43} as

$$D^\circ(\text{M-X}) = \Delta H_f^\circ(\text{M}) + \Delta H_f^\circ(\text{X}) - \Delta H_f^\circ(\text{MX}). \quad (2)$$

The values of ΔH_f° at 298.15 K were referred from an earlier report of the literature.⁴⁴ The ΔH_f° values used for calculating the bond energy (D°) via thermochemical cycles are presented in Table S1 of the supplementary material. Table 1 presents the calculated surface energies of (100) and (111) planes of Cu₃N and some ionic metal oxides. Surface energies obtained from this calculation tend to be greater than those obtained from DFT calculations reported earlier. The surface energies of the (111) plane in Cu₃N and MgO with cubic crystal structures tend to be larger than those of the (100) plane. The (111) surface energies of various rock salt type crystal structure compounds such as

MgO(111) and Cu₃N tend to be sufficiently larger than those of (100) plane surface energies, indicating that the (111) surfaces of rock salt type compounds and Cu₃N are regarded as having lower stability than their (100) surfaces,^{45,46,50} which are consistent with the trend of this calculation. The calculated surface energy of the (111) plane of Cu₃N is higher than that of the Cu₃N(100) surface, but both the (100)-orientation and (111)-orientation of Cu₃N are generated during thin film growth, as depicted in Fig. 1. This finding implies that the (111)-Cu₃N orientation probability is approximately equal to that of the (100)-Cu₃N orientation probability. During the thin film growth of Cu₃N using the reactive DC magnetron sputtering method, the (111)-oriented Cu₃N growth was promoted by overcoming the influence of high surface energy. Table 1 also presents the surface energies of the substrates used for this study.

To investigate the influence of a substrate on the (111)-orientation and (100)-orientation in Cu₃N layers, XRD phi (ϕ) scan measurements were taken. Figure 2 shows the XRD ϕ scan patterns of the majority orientations of Cu₃N layers grown on the MgO(100) and MgO(111) substrates. As depicted in Figs. 1(a) and 1(d), the majority orientations of Cu₃N layers are (100) of Cu₃N for the MgO(100) substrate, (111) of Cu₃N for the MgO(111) substrate. The second-largest XRD intensity orientations are the (111)-orientation of Cu₃N for the MgO(100) substrate, and the (100)-orientation of Cu₃N for the MgO(111) substrate. For the second-largest XRD intensity orientations, taking XRD ϕ scan measurements was difficult because of their weak XRD peak intensities and broadening along the ψ direction, indicating the large mosaicities of these domain rotations, as portrayed in the XRD 2D images of Fig. 1. The XRD ϕ scan patterns in Fig. 2 were obtained from the same samples as those shown in Figs. 1(a)(b) and 1(g)(h). In the Cu₃N layer grown on the MgO(100) substrate, fourfold symmetric peaks of Cu₃N{111} diffraction were observed (Fig. 2(a)). By contrast, in the Cu₃N layer grown on the MgO(111) substrate, threefold symmetric peaks of Cu₃N{111} were observed (Fig. 2(c)). Bump-like weak peaks observed between the neighboring sharp peaks in Fig. 2(c) originate from the broadened reciprocal lattice point because of the large mosaicity along the ψ direction of the second-largest XRD intensity orientations. The in-plane crystallographic epitaxial

relations of the majority orientations are (100)Cu₃N//(100)MgO and [010]Cu₃N//[010]MgO for the Cu₃N thin film on the MgO(100) substrate and (111)Cu₃N//(111)MgO and [1 $\bar{1}$ 0]Cu₃N//[1 $\bar{1}$ 0]MgO for the Cu₃N thin film on the MgO(111) substrate, which indicates that the Cu₃N majority orientations were grown under the strong influence of the MgO substrate surface lattices. However, the second-largest XRD intensity orientations are regarded as growing under conditions of lesser influence of the MgO substrate surface lattices because of the large mosaicities of the domain rotation along the ψ directions (Figs. 1(a) and 1(g)). The lattice mismatch values between Cu₃N (a cubic anti-ReO₃ structure, $a = 3.819 \text{ \AA}$) and MgO (a cubic rock salt crystal structure, $a = 4.213 \text{ \AA}$) substrates are approximately 9.35% for both between Cu₃N(100) and MgO(100), and between Cu₃N(111) and MgO(111). On MgO(100) substrates, the second-largest XRD intensity orientations of Cu₃N(111) have a surface unit cell with rhombohedral symmetry, whereas the surface unit cell of the MgO(100) substrate has cubic symmetry. Similarly, on MgO(111) substrates, the surface unit cell of the second-largest XRD intensity orientations for Cu₃N(100) has cubic symmetry, whereas the surface unit cell of the MgO(111) substrate has rhombohedral symmetry. Because the surface unit cell symmetries of the second-largest XRD intensity orientations and MgO substrates mutually differ, the second-largest XRD intensity orientations are regarded as growing irrespective of the substrate surface lattice influences. Although the majority orientations grew under the influence of the substrate surface lattices, both the second-largest XRD intensity (100)-orientation and (111)-orientation of Cu₃N grew despite their surface energies and the MgO substrate surface lattice symmetries.

To control epitaxial thin film growth, the effects of thermodynamics and kinetics should be considered. To this end, the growth temperature and growth rate effects are evaluated using analyses of XRD 2θ - ω patterns. All the XRD peak intensities originating from Cu₃N layers are normalized by both the film thickness d and substrate XRD peak intensities, i.e., $I_{norm}^{(111)} = I_{film}^{(111)} / (I_{substrate} d)$.^{51,52} The following equation was used to evaluate the intensity ratio of the summation of Cu₃N(111) and (222) peaks to the sum of the XRD peak intensities of all Cu₃N.

$$I_R^{(111)+(222)} = \frac{I_{norm}^{(111)} + I_{norm}^{(222)}}{\Sigma I_{norm}^{(hkl)}} \quad (3)$$

Therein, $I_{norm}^{(111)}$ and $I_{norm}^{(222)}$ respectively represent the normalized XRD intensity peaks of $\text{Cu}_3\text{N}(111)$ and (222) . In addition, $\Sigma I_{norm}^{(hkl)}$ represents the summation of intensities of all Cu_3N XRD peaks. The ratio is represented by $I_R^{(111)+(222)}$. To elucidate the influence of growth parameters on the crystallographic orientation of Cu_3N thin films, $I_R^{(111)+(222)}$ as a function of the growth temperature and growth rate is discussed. In this case, the XRD $2\theta-\omega$ patterns of 52 samples obtained on $\text{MgO}(100)$, $\text{SrTiO}_3(100)$, sapphire(0001), and $\text{MgO}(111)$ substrates are used. The results demonstrate that $I_R^{(111)+(222)}$ depends on the growth temperature and growth rate, as shown in Figs. 3 and 4. In addition, $I_R^{(111)+(222)}$ tends to decrease as the growth temperature increases. Moreover, it exhibits a maximum value at a growth rate of 0.33 \AA/s . Below that growth rate of 0.33 \AA/s , $I_R^{(111)+(222)}$ tends to decrease. As shown in Fig. S3, the growth rate showed a tendency to remain largely unchanged at growth temperatures of $110 \text{ }^\circ\text{C}$ to $200 \text{ }^\circ\text{C}$ at the constant of other growth parameters. Orientation-pure, single (111)-oriented Cu_3N epitaxial thin films have been obtained at the low growth temperature of $110 \text{ }^\circ\text{C}$ and low growth rate of 0.33 \AA/s ($I_R^{(111)+(222)} = 1$). This growth condition is an appropriate growth condition to achieve a higher formation probability of Cu_3N (111)-orientation possessing large surface energy, in which the balance between thermodynamics and kinetics is tuned.

By contrast, the behaviors of $I_R^{(100)+(200)}$ for the growth temperature and the growth rate differ from those of $I_R^{(111)+(222)}$ as shown in Figs. 3 and 4. Even though the surface energy of $\text{Cu}_3\text{N}(100)$ is less than that of $\text{Cu}_3\text{N}(111)$, the values of $I_R^{(100)+(200)}$ are low, with a maximum value of 0.67. The variations in the values of $I_R^{(100)+(200)}$ to the growth temperature and the growth rate tend to be small compared to those of $I_R^{(111)+(222)}$. This result indicates that, in the reactive DC magnetron sputtering, the formation probability of $\text{Cu}_3\text{N}(111)$ orientation with larger surface energy

remains as well as that of $\text{Cu}_3\text{N}(100)$ orientation with smaller surface energy. Actually, $\text{SrTiO}_3(100)$ ($a = 3.905 \text{ \AA}^4$) substrate is a promising substrate for $\text{Cu}_3\text{N}(100)$ growth because of its small lattice mismatch with $\text{Cu}_3\text{N}(100)$ of 2.2%. However, the $I_R^{(100)+(200)}$ values found from this study are low even with the $\text{SrTiO}_3(100)$ substrate. The formation of crystal nuclei with high surface energy orientations is regarded as suppressed at the initial thin film growth stage because of their instability. However, with Cu_3N thin film growth via reactive DC magnetron sputtering, high surface energy crystal orientations are formed. This study presents an example that overcomes the relation between the preference for out-of-orientation in thin film growth and the stability of the resultant surface. Additional effects on Cu_3N thin film growth are surface reconstruction effects of Cu_3N and substrates, along with non-charge-neutrality effects of $\text{Cu}_3\text{N}(111)$ and $\text{MgO}(111)$ substrate surfaces. Further detailed research must be undertaken to elucidate these effects on Cu_3N thin film growth.

For this study, pure (111)-oriented Cu_3N epitaxial thin films were grown on $\text{MgO}(111)$ substrates with a growth temperature of $110 \text{ }^\circ\text{C}$ and a growth rate of 0.33 \AA/s by adjusting the balance between thermodynamics and kinetics. A strong and sharp (111) diffraction peak from the Cu_3N thin film is visible, aligned with the (111) peak of the MgO substrate, as shown in Figs. 5(a) and 5(b). The absence (or marked suppression) of other Cu_3N orientations in the XRD diffraction pattern supports the epitaxial characteristics of the growth. The lattice constant obtained for the Cu_3N thin film is 3.814 \AA with 0.12% expansion, compared to the bulk lattice constant.⁹ The fairly good lattice mismatch (9.35%) and surface atomic configuration symmetry between $\text{Cu}_3\text{N}(111)$ and $\text{MgO}(111)$ facilitate this orientation. Single orientation along the in-plane direction was also found from XRD ϕ scanning measurements, as shown in Figs. 5(c) and 5(d), corresponding to the in-plane epitaxial relation of $\text{Cu}_3\text{N}[\bar{1}\bar{1}0]//\text{MgO}[\bar{1}\bar{1}0]$. The XRD intensity of $\text{Cu}_3\text{N}\{111\}$ peaks in the ϕ scan pattern for the single (111)-oriented epitaxial Cu_3N thin film is much greater than the intensity found for the polycrystalline Cu_3N thin film grown on a $\text{MgO}(111)$ substrate, as presented in Fig. 2(c). The noise level in the XRD ϕ scan pattern of the single (111)-oriented epitaxial Cu_3N thin film (Fig. 5(c)) is much lower than that of the polycrystalline Cu_3N thin film presented in Fig. 2(c).

By contrast, single (100)-oriented Cu_3N epitaxial thin films were fabricated on $\text{SrTiO}_3(100)$ substrates using molecular beam epitaxy.¹⁸ Generally speaking, the energy range of the supply species used with sputtering methods (order of 10 eV)²⁸ is much greater than that of supply species used with molecular beam epitaxy methods (thermal energy on the order of 0.1 eV).¹ Therefore, thin films produced using molecular beam epitaxy are grown under conditions closely approximating thermodynamic conditions, whereas thin films produced by sputtering are grown under conditions that differ greatly from thermodynamic conditions. In addition to the energy of supply species, with molecular beam epitaxy, supply species reach substrate surfaces by maintaining their mean free path in sufficiently high-vacuum conditions. Consequently, the supply species are Cu atom species and N atom species. By contrast, with reactive DC magnetron sputtering, the supply species might be Cu_xN_y in addition to Cu-related and N-related species with higher energies under high N_2 gas pressures on the order of several pascals.⁵³ Differences in supply species might also influence the crystal orientation control. In addition to the differences in supply species and their energies, the surface energies on crystal growth differ according to their growth conditions and growth methods.^{2,4} Therefore, for the reactive DC magnetron sputtering thin film growth, it can be inferred that the formation probability of (111)-orientation of Cu_3N with large surface energy exists in addition to the (100)-orientation formation probability, irrespective of substrate surface unit cell symmetry.

Optical absorption spectra of the single (111)-oriented Cu_3N epitaxial film grown on a $\text{MgO}(111)$ substrate, as obtained by spectroscopic ellipsometry measurements, are presented in Fig. 6(a). The direct band gap was obtained by plotting the $(\alpha\hbar\omega)^2$ versus $\hbar\omega$ (energy) curve and by extrapolating the full line to the horizontal axis of $\hbar\omega$. Similarly, the indirect band gap was found by extrapolating the full line to the horizontal line of $\hbar\omega$ in the $(\alpha\hbar\omega)^{1/2}$ versus $\hbar\omega$ curve. The direct and indirect band gap values were found respectively as 1.80 and 0.82 eV. Various values for the optical band gap of Cu_3N have been reported to date. Experimentally obtained values for optical band gaps have been reported as 1.3–1.7 eV,¹⁶ 1.44 eV,²⁴ 1.5–1.8 eV,¹⁹ 1.57 eV,¹⁵ 1.65 eV,¹² 1.68 eV,²¹ 1.7 eV,²² 1.7–1.9 eV,²⁵ 1.72–2.38 eV,²⁰ 1.8 eV,¹¹ 2.14 eV, and 2.21 eV²⁶ for direct optical band gaps, and as

1.0 eV,²² 1.17–1.68 eV,²⁰ 1.3 eV,¹⁸ 1.45 eV, and 1.46 eV²⁶ for indirect optical band gaps. The values of the direct and indirect optical band gaps obtained from this study are reasonably consistent with those reported from most of the earlier studies conducted on this subject. The direct and indirect optical band gaps of Cu₃N obtained from this study are discussed by comparing them with the electronic band structure calculated using DFT, as presented in Fig. 6(b). The electronic band structure in Fig. 6(b) shows similar trends to those reported earlier.^{21,24,54,55} The direct transition band gap is apparent between the R points, with a value of 1.21 eV. The electronic structure was calculated using mBJ, which provides improved band gap values compared to DFT calculations using LDA and GGA-PBE. However, even the band gap values obtained using mBJ calculations generally tend to show smaller band gaps than the experimentally obtained values.^{34–36} Therefore, this relation is consistent with the direct optical band gap of 1.80 eV obtained from this study. In Fig. 6(b), the indirect band gaps between the Γ (conduction band minimum, CBM)–R (valence band maximum, VBM) points and between the M (CBM)–R (VBM) points are apparent, respectively, with values of 1.08 eV and 0.41 eV. The indirect optical band gap of 0.82 eV observed in Fig. 6(a) is regarded as attributable to these indirect transitions in Fig. 6(b). The two broad peaks observed at the low energy side of the $(\alpha\hbar\omega)^{1/2}$ versus the $\hbar\omega$ curve in Fig. 6(a) were also reported earlier in the literature.^{21,24,26} The origin of these absorption peaks remains unclear. One possible explanation is absorption because of the energy levels formed within the band gap, caused by factors such as defects or impurities.^{22,55–57} More-detailed studies must be conducted to elucidate these two broad absorption peaks. The optical bandgap for direct transition is larger than the value obtained by DFT calculation. In contrast, the optical bandgap for indirect transition was similar to the values calculated using DFT. One factor that might be related to this point is the influence of the functionals on DFT calculations. On the lowest conduction band, the energy level at the Γ point for mBJ is lower than that at the R point, whereas the energy level at the Γ point is higher than that at the R point for GGA-PBE, as presented in Fig. S5 and Table S2. Changes in band structures attributable to functional differences have been reported,^{58–63} as described in detail in the supplementary material. Comparison of the

optical absorption spectra with the electronic band structure calculated using DFT indicates that the single (111)-oriented Cu₃N epitaxial films grown on MgO(111) substrates form an appropriate electronic structure, thereby indicating the high-crystalline quality epitaxial thin films.

4. CONCLUSION

In conclusion, thin film growth of Cu₃N by reactive DC magnetron sputtering was conducted on MgO(100), SrTiO₃(100), sapphire(0001), and MgO(111) substrates. Irrespective of the substrate, a tendency to form polycrystalline orientations of Cu₃N phases was observed. The dominant crystalline orientations of Cu₃N were (111) and (100) orientations, although the surface energy of Cu₃N(111) is higher than that of Cu₃N(100). By tuning the parameters of thermodynamics and kinetics affecting thin film growth, single (111)-oriented Cu₃N epitaxial thin films were obtained on MgO(111) substrates, which possess high surface energy. Optical direct-band and indirect-band gaps of 1.80 eV and 0.82 eV in the single (111)-oriented Cu₃N epitaxial thin film were observed based on spectroscopic ellipsometry measurements, which are reasonably consistent with values presented in most earlier reports. This study has highlighted the importance of the following points. For thin film growth, even a crystal orientation with high surface energy can be grown stably because the surface energy varies with growth conditions and growth methods, in addition to temperatures, absorption, and other parameters. Such single-oriented epitaxial thin films are producible by tuning the balance between thermodynamics and kinetics during thin film growth and by selecting an appropriate substrate.

SUPPLEMENTARY MATERIAL

See the supplementary material for AES spectra, AFM image, thermochemical cycles and formation enthalpies for calculating the bond energy, growth rate of Cu₃N thin films shown as a function of growth temperature, XRD 2 θ - ω patterns of a Cu₃N(111) epitaxial thin film grown on MgO(111) substrate, and of a bare MgO(111) substrate, electronic band structures of Cu₃N calculated by mBJ exchange potential and GGA-PBE exchange-correlation functional, and raw data of spectroscopic ellipsometry measurements (amplitude ratio (Ψ), phase change (Δ), complex refractive indices (ϵ_1 and ϵ_2), refractive index (n), extinction coefficient (k), and absorption coefficient (α)).

ACKNOWLEDGMENTS

This work was supported by funding from JSPS-KAKENHI (no. 23H00263) and the JST–Mirai Program (JPMJMI19A1). We thank Ms. Tomoko Kobayashi for AFM measurements.

AUTHOR DECLARATIONS

Conflict of Interest

The authors have no conflict to disclose.

Author Contributions

Zainab Fatima: Investigation (equal); Data curation (equal); Formal analysis (equal); Writing/original draft preparation (equal). **Isao Ohkubo:** Conceptualization (equal); Methodology (equal); Investigation (equal); Data curation (equal); Formal analysis (equal); Writing/original draft preparation (equal); Project administration (equal); Funding acquisition (equal). **Satoshi Ishii:** Methodology (equal); Resources (equal); Investigation (equal); Data curation (equal); Formal analysis (equal); Writing/review & editing (supporting). **Takahiro Nagata:** Methodology (equal); Resources (equal); Investigation (equal); Data curation (equal); Formal analysis (equal); Writing/review & editing (supporting). **Takashi Aizawa:** Conceptualization (equal); Methodology (equal); Investigation (equal); Writing/review & editing (equal). **Takao Mori:** Resources (equal); Writing/review & editing (supporting); Supervision (equal); Project administration (equal); Funding acquisition (equal).

DATA AVAILABILITY

Data supporting the findings of this study are available within the article and its supplementary material, and upon reasonable request from the corresponding author.

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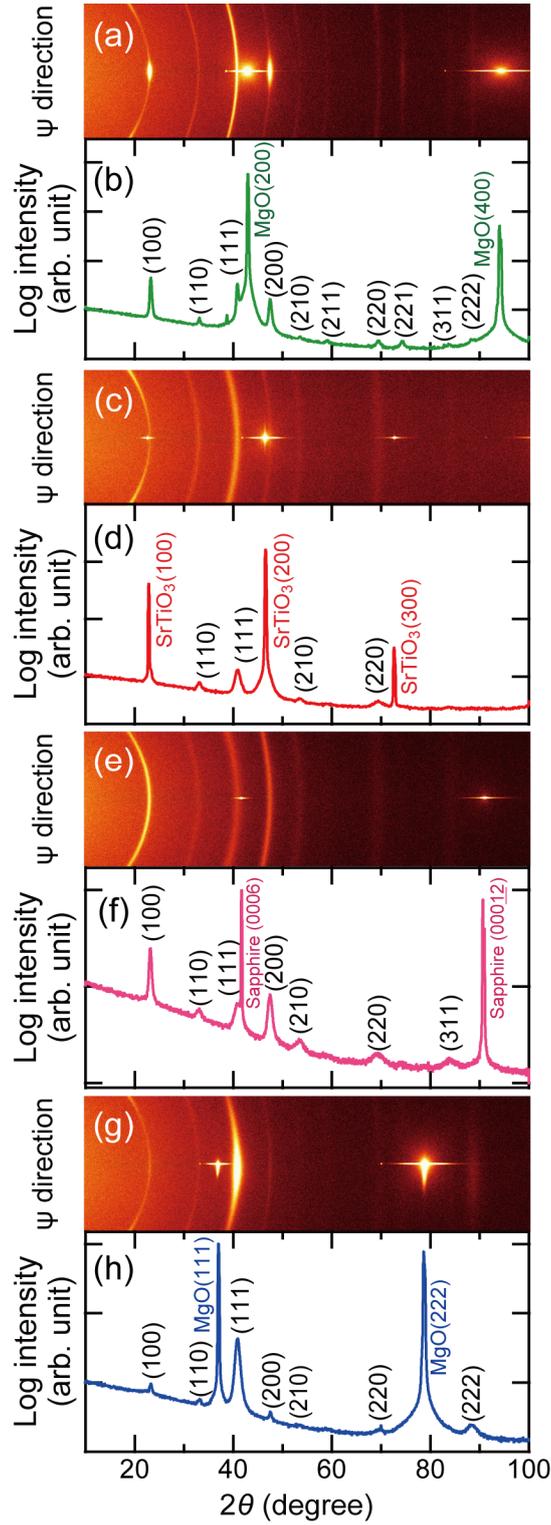


Fig. 1. (a),(c),(e),(g) XRD 2D images and (b),(d),(f),(h) 2θ - ω patterns of Cu_3N thin films grown with the following conditions and substrate: (a),(b) $\text{MgO}(100)$ substrate at 170°C growth temperature, DC power 25 W, nitrogen gas partial pressure (P_{N_2}) 1.7 Pa; (c),(d) $\text{SrTiO}_3(100)$ substrate at 180°C , 25 W, 1.7 Pa; (e),(f) sapphire(0001) substrate at 170°C , 40 W, 1.7 Pa, and (g),(h) $\text{MgO}(111)$ substrate at 180°C , 25 W, 1.8 Pa. The vertical axis of XRD 2D images corresponds to the ψ direction.

Table 1. Surface energy values and parameters used for surface energy calculation: number of broken bonds per surface unit cell ($S_{(hkl)}$), coordination number of metal cation (CN), bond energy (D°), surface energy $\gamma_{(hkl)}$, and surface energy calculated using density-functional theory calculation (DFT) referred from the literature.

CF	Orientation	$S_{(hkl)}$	CN	D° (kJ/mol)	$\gamma_{(hkl)}$ (this study) (J/m ²)	$\gamma_{(hkl)}$ (DFT) (J/m ²)
Cu ₃ N	(100)	1 ^a			2.69	0.123, 2.771 ^b
Cu ₃ N	(111)	3 ^a	2	235.53	4.64	0.394 ^b
MgO	(100)	2 ^a			6.23	0.92–1.03 ^c
MgO	(111)	3 ^a	6	166.43	10.79	1.37 ^c
Sapphire (α -Al ₂ O ₃)	(0001)	6 ^a	6	256.33	13.02	2.04, 3.99 ^d
SrTiO ₃	(100)	4 (Sr-O)	12 (Sr)	167.60 (Sr-O)	10.77	1.21, 1.29 ^f
		1 (Ti-O)	6 (Ti)	318.82 (Ti-O)		

^a The numbers of broken bonds of the Cu plane and the Cu-N plane of Cu₃N(100), the Cu plane and the N plane of Cu₃N(111), the Mg plane and the O plane of MgO(111), the Al plane and the O plane of sapphire(0001), and the SrO plane and the TiO₂ plane of SrTiO₃(100) are, respectively, equal. ^b Ref. 45, ^c Ref. 46, ^d Ref. 47, ^e Ref. 48, ^f Ref. 49.

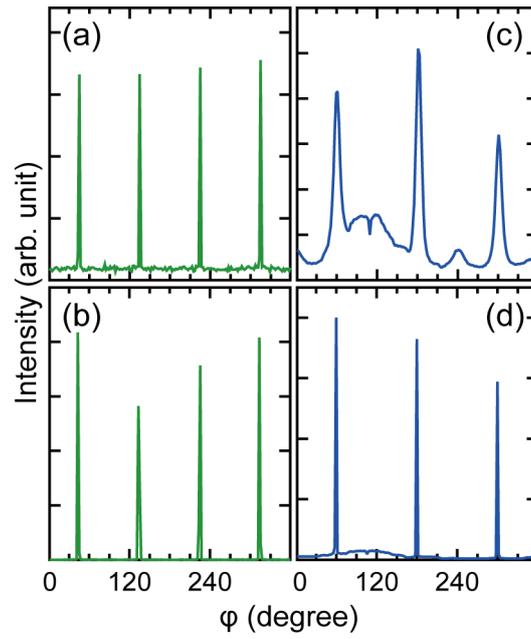


Fig. 2. XRD ϕ (ϕ) scan patterns of (a),(b) a Cu_3N thin film grown on a $\text{MgO}(100)$ substrate and (c),(d) a Cu_3N thin film grown on a $\text{MgO}(111)$ substrate. The measured diffraction planes are (a),(c) $\text{Cu}_3\text{N}\{111\}$ and (b),(d) $\text{MgO}\{111\}$ for both on a $\text{MgO}(100)$ substrate and on a $\text{MgO}(111)$ substrate.

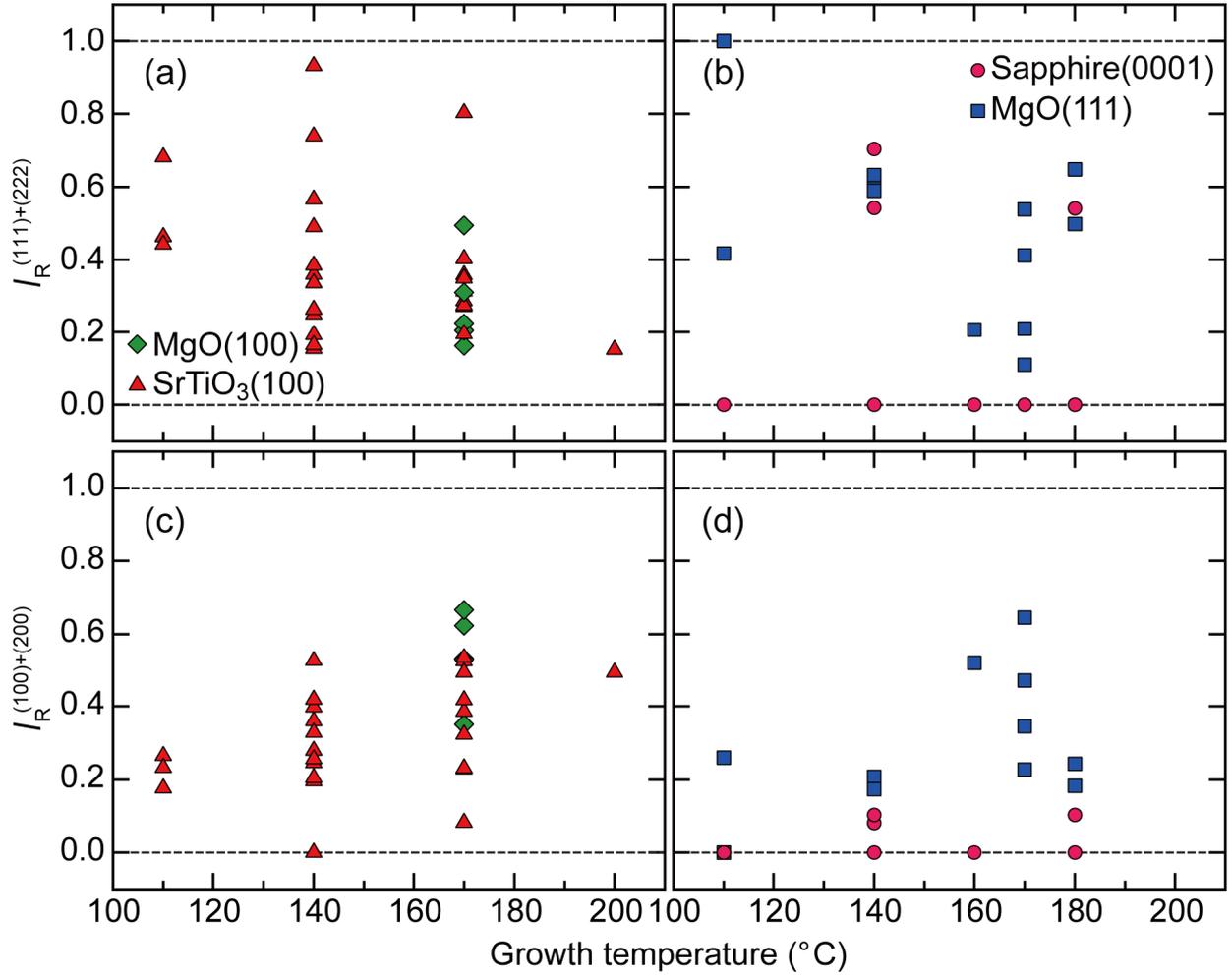


Fig. 3. (a),(b) Ratio ($I_R^{(111)+(222)}$) of the normalized XRD intensity $I_{norm}^{(111)}$ and $I_{norm}^{(222)}$ with $\Sigma I_{norm}^{(hkl)}$ shown as a function of growth temperature: (a) cubic symmetry surface unit cell substrates of MgO(100) and SrTiO₃(100); (b) rhombohedral symmetry surface unit cell substrates of sapphire(0001) and MgO(111). (c),(d) Ratio ($I_R^{(100)+(200)}$) of normalized XRD intensity $I_{norm}^{(100)}$ and $I_{norm}^{(200)}$ with $\Sigma I_{norm}^{(hkl)}$ shown as a function of growth temperature: (c) cubic symmetry surface unit cell substrates of MgO(100) and SrTiO₃(100); and (d) rhombohedral symmetry surface unit cell substrates of sapphire(0001) and MgO(111).

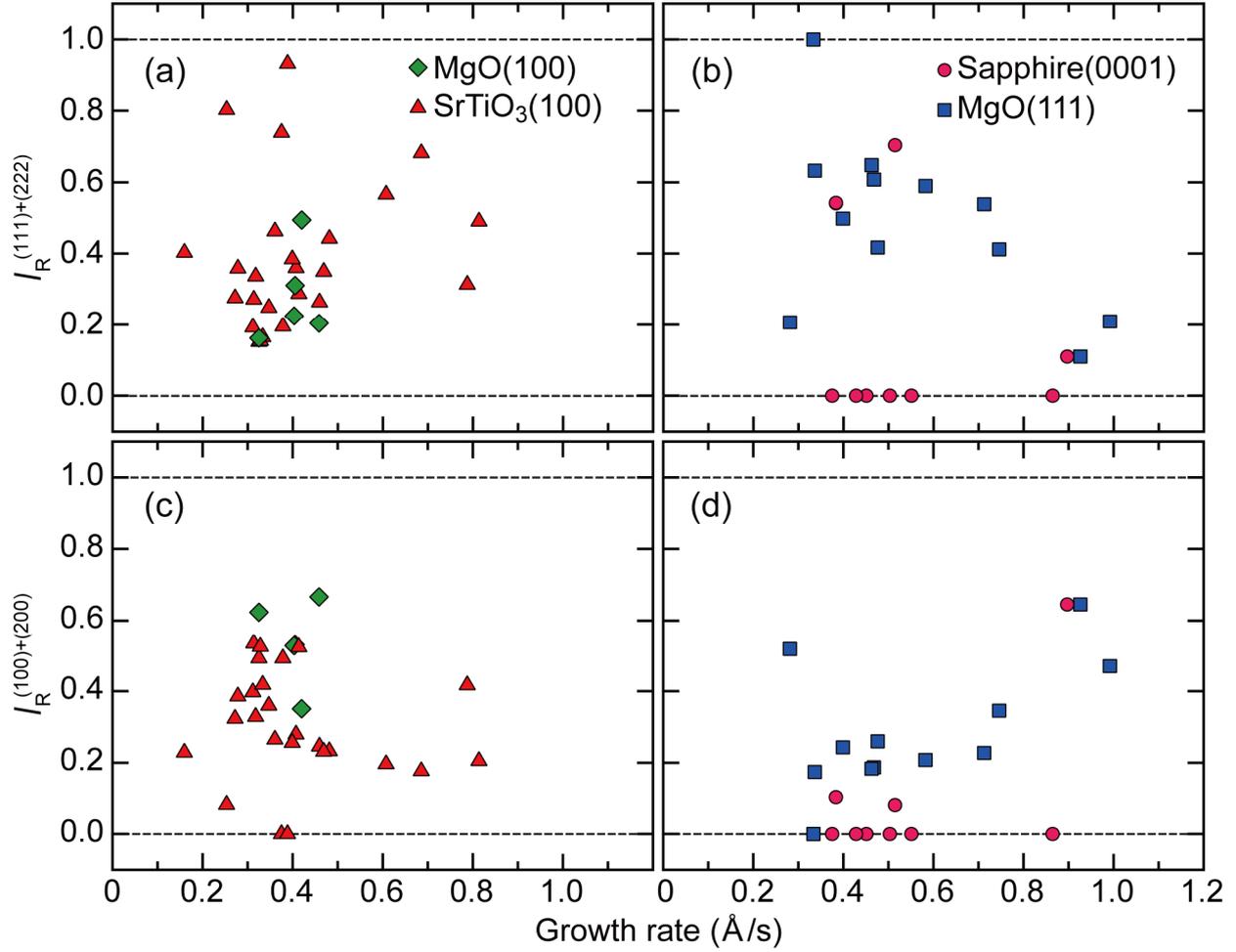


Fig. 4. (a),(b) Ratio ($I_R^{(111)+(222)}$) of the normalized XRD intensity $I_{norm}^{(111)}$ and $I_{norm}^{(222)}$ with $\Sigma I_{norm}^{(hkl)}$ shown as a function of growth rate: (a) cubic symmetry surface unit cell substrates of MgO(100) and SrTiO₃(100); (b) rhombohedral symmetry surface unit cell substrates of sapphire(0001) and MgO(111). (c),(d) Ratio ($I_R^{(100)+(200)}$) of normalized XRD intensity $I_{norm}^{(100)}$ and $I_{norm}^{(200)}$ with $\Sigma I_{norm}^{(hkl)}$ shown as a function of growth rate: (c) cubic symmetry surface unit cell substrates of MgO(100) and SrTiO₃(100); and (d) rhombohedral symmetry surface unit cell substrates of sapphire(0001) and MgO(111).

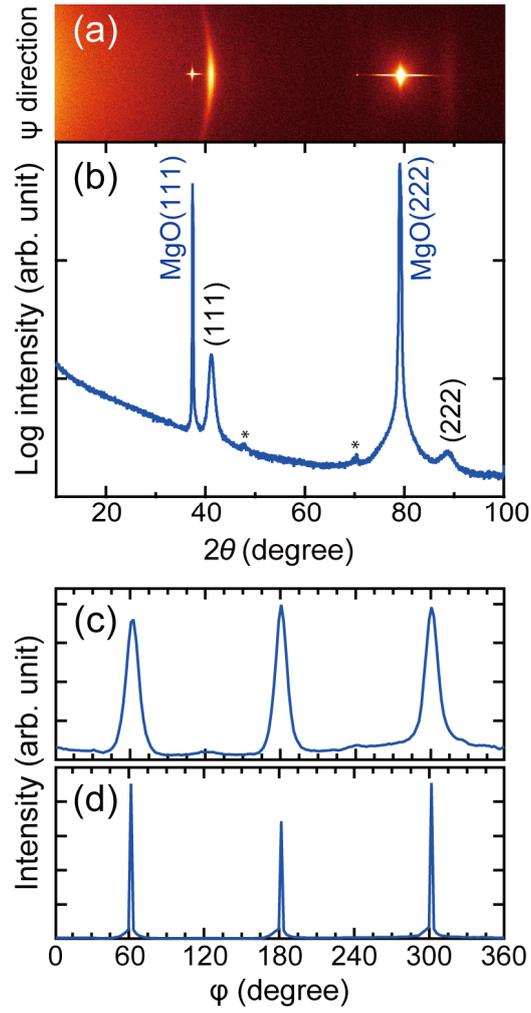


Fig. 5. (a) XRD 2D image, (b) 2θ - ω pattern, (c),(d) ϕ scan patterns of a single (111)-oriented Cu_3N epitaxial thin film grown on a $\text{MgO}(111)$ substrate. The vertical axis of the XRD 2D image corresponds to the ψ direction. The asterisk * in Fig. 5 (b) denotes peak origination from the $\text{MgO}(111)$ substrate, as shown in Fig. S4. The measured diffraction planes for ϕ scans are (c) $\text{Cu}_3\text{N}\{111\}$ and (d) $\text{MgO}\{111\}$.

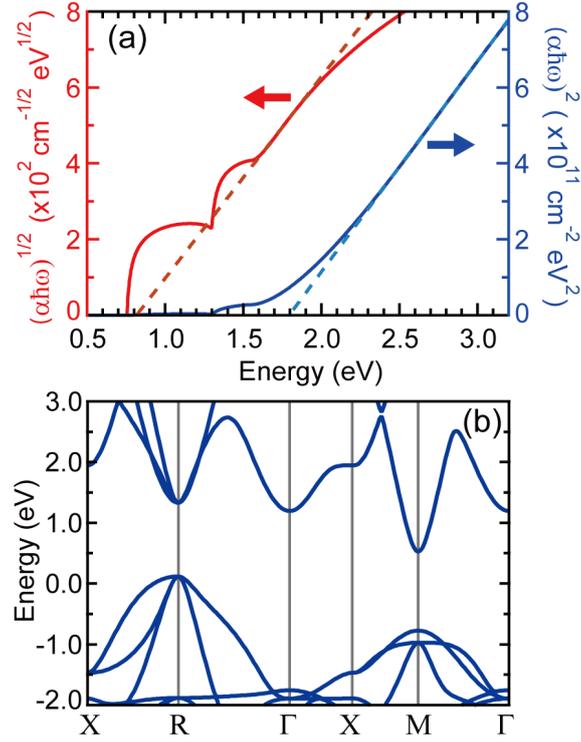


Fig. 6. (a) Optical absorption spectra of a single (111)-oriented Cu₃N epitaxial film grown on a MgO(111) substrate and (b) an electronic band structure of Cu₃N obtained using density-functional theory calculation using the mBJ exchange potential. In the optical absorption spectra, the $(\alpha\hbar\omega)^{1/2}$ versus $\hbar\omega$ plot is used for determination of the indirect energy band gap; also, the $(\alpha\hbar\omega)^2$ versus $\hbar\omega$ plot is used for calculation of the direct energy band gap. The labels in the Brillouin zone of electronic band structure Γ , X, R, and M respectively correspond to the following in lattice constant units: 0,0,0; 1,0,0; 1,1,1; and 1,0,1.